

Introduction to Inertial Confinement Fusion (ICF)

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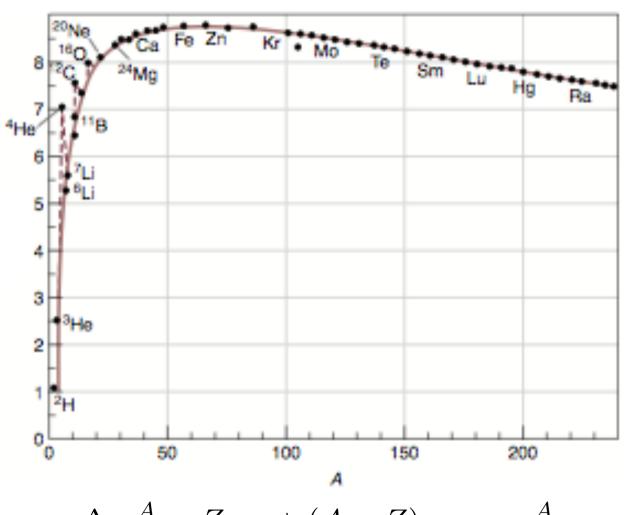
The voyage of nuclear fusion has started about 60 years ago (Sacharov, Teller, ...) and despite many progress has mainly provided disillusions...

50 years ago the laser was invented, opening the field of "Inertial Fusion" (Basov, Nuckolls, ...)

Today we are probably close to the demonstration of ignition, the scientific feasibility of fusion, which will conclude the first part of this travel.



Weizsaker semiempirical mass formula



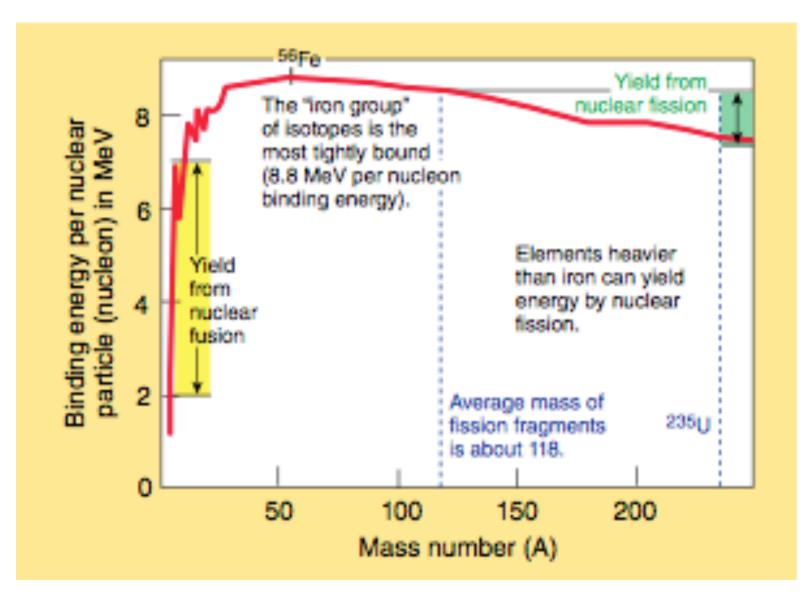
$$B_Z^A = \Delta m_Z^A c^2 / A$$

Fig. II-IO The binding energy per nucleon versus atomic mass number A. The solid curve represents the Weizsäcker semiempirical binding-energy formula, Equation II-I2.

 $\Delta m_Z^A = Zm_p + (A - Z)m_n - m_Z^A.$

Weizsaker semiempirical mass formula



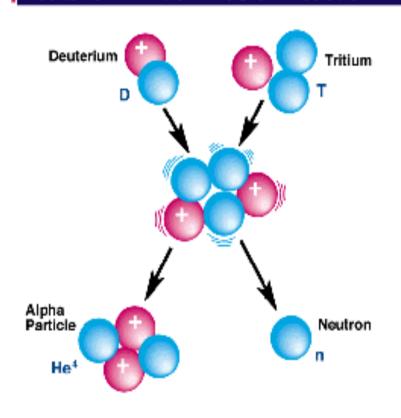


 $E=mc^2$



Thermonuclear Fusion:

Deuterium-Tritium Fusion Reaction



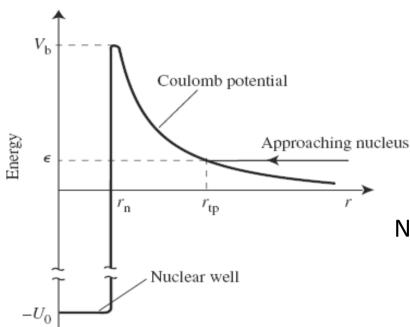
$$D + T \rightarrow He^4 + n + 17,6 MeV$$

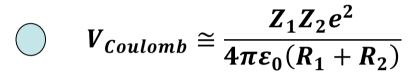
1/5 (17.6 MeV) = 3.5 MeV for the α particle

4/5 (17.6 MeV) = 14.1 MeV for the neutron



Coulomb repulsion:





Nuclear radius depends on A as

$$R_i \cong 1.4 \ 10^{-15} A_i^{1/3}$$

The energy threshold ≈ 400 KeV, i.e. $T \approx 4.6$ 10^9 K (1 eV = 11400 K). For comparison $T \approx$ $1.6 \cdot 10^7$ K at sun core.

→ Tunnel effect

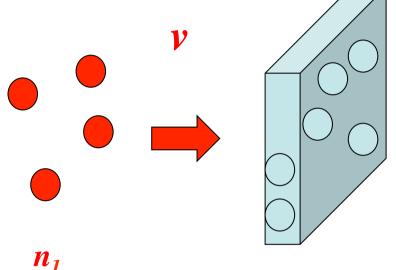


Reaction probability and cross section:

$$\frac{dn}{dt} = \sigma(v)v n_1 n_2$$

 n_1 , n_2 densities of particles of species 1 and 2 (#/cm³)

dn/dt number of reaction per unit volume and unit time

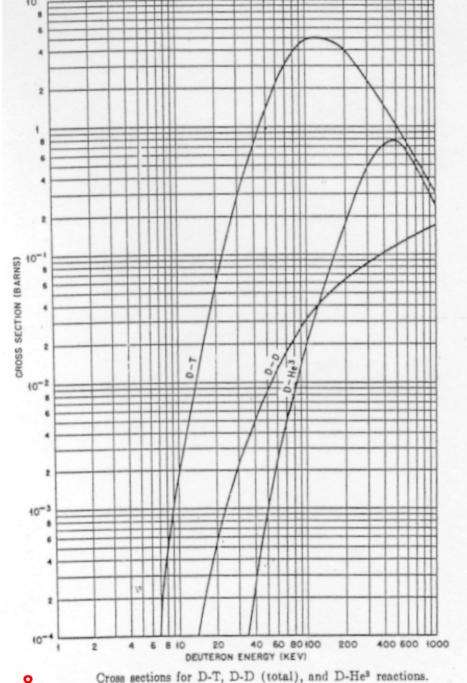


 n_2

 $n_1 v = \#/\text{cm}^2 \text{sec} = \text{flux of incident particles}$

 σ has the units of surface (cm²)

 $n_1 \sigma$ = total "covered" surface per unit volume.





D + D
$$\rightarrow$$
 T + H¹ + 4,03 MeV
D + D \rightarrow He³ + n + 3,27 MeV

$$D + He^3 \rightarrow He^4 + H^1 + 18,3 MeV$$

$$D + T \rightarrow He^4 + n + 17,6 MeV$$

D-T fusion reaction has a larger probability and peaks at lower energy

Reaction rate:

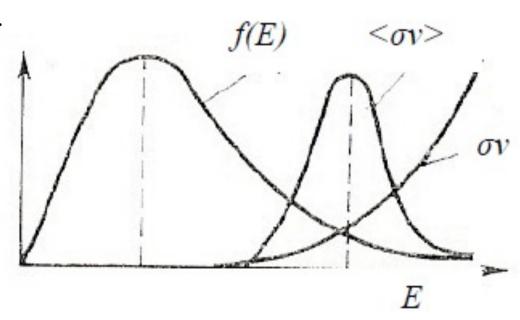


In a plasma we must average the cross section over the distribution of velocities

$$f_M(v) = (2\pi)^{-3/2} \exp(-m_r v^2/2T)$$

$$\langle \sigma v \rangle = 4\pi \int_0^\infty v^3 \sigma(v) f_M(v) dv.$$

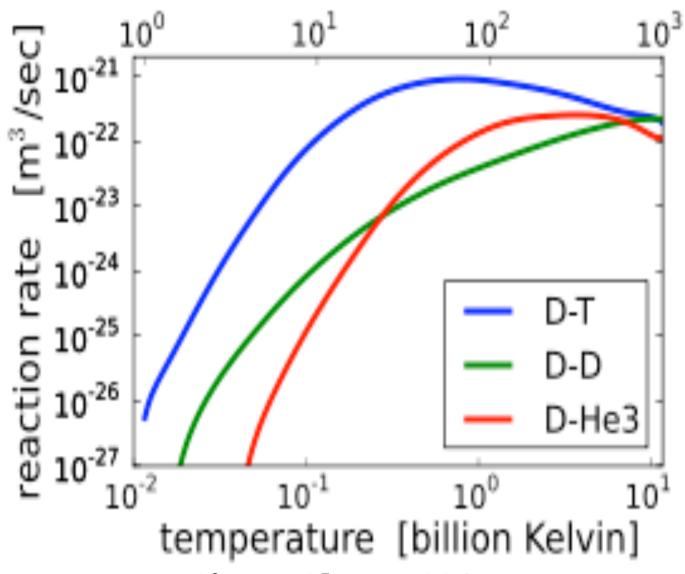
At low temperature the main contribution to fusion reactions comes from ions in the distribution tail



D.Batani, 1. Introduction to Fusion, 2016

temperature [keV]





Fusion needs high temperatures (T> 5 keV)

 $10^9 \text{ K} \approx 10^5 \text{ eV} = 100 \text{ keV}$

Reaction rate:



DT reaction

$$\langle \sigma_{DT} v \rangle \simeq 9.1 \times 10^{-16} \exp\left(-0.572 \left| \ln(T_{\text{keV}}/64.2) \right|^{2.13}\right) \text{ cm}^3/\text{s.}$$

$$\langle \sigma_{DT} v \rangle \simeq C_{DT}^b T^3 = 1.1 \times 10^{-19} T_{\text{keV}}^3 \text{ cm}^3/\text{s.}$$

$$3-100 \text{ keV}$$

$$3-8 \text{ keV}$$

$$\langle \sigma_{DT} v \rangle \simeq C_{DT} T^2 = 1.1 \times 10^{-18} T_{\text{keV}}^2 \text{ cm}^3/\text{s.}$$

$$8-20 \text{ keV}$$

DD reaction

$$\langle \sigma_{DDn} v \rangle \simeq 2.7 \times 10^{-14} T_{\text{keV}}^{-2/3} \exp\left(-19.8 T_{\text{keV}}^{-1/3}\right) \text{ cm}^3/\text{s.}$$
3-50 keV



Plasma conditions for fusion: confinament time

Confinament time:
$$\tau = \frac{W}{P_{losses}}$$

W = internal energy per unit volume (J/m^3)

 P_{losses} = dissipated power per unit volume (this is mainly radiation losses due to bremsstrahlung).

 τ = the time over which the system is able to keep its energy)

$$W = \frac{3}{2} \left((n_e + n_i)kT \right) = 3n_e T(eV)$$

Rate of energy production

$$P_{fusion} = n_D n_T \langle \sigma v \rangle E_\alpha = 4n_e^2 \langle \sigma v \rangle E_\alpha$$



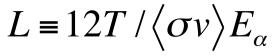
LAWSON'S CRITERIUM

$$P_{fusion} > P_{losses} \qquad \Rightarrow \frac{1}{4} n_e^2 \langle \sigma v \rangle E_\alpha > 3n_e T / \tau$$

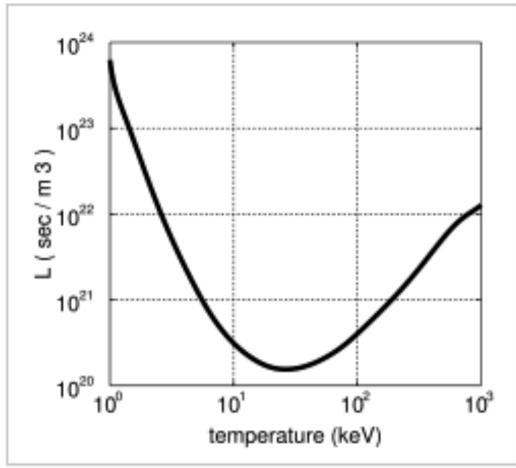
$$\Rightarrow n_e \tau > 12T / \langle \sigma v \rangle E_\alpha \equiv L$$

 $<\sigma v>$ depends on T.

[
$$E_{\alpha} = 3.5 \text{ MeV}$$
]







Is function of T and has a minimum at $\approx 25 \text{ keV}$ (2.910⁸ K) for D + T \rightarrow ⁴He + n

The deuterium-tritium L function (minimum n_eT_E = needed to satisfy the Lawson criterion) minimizes near the temperature 25 keV (300 million kelvins).

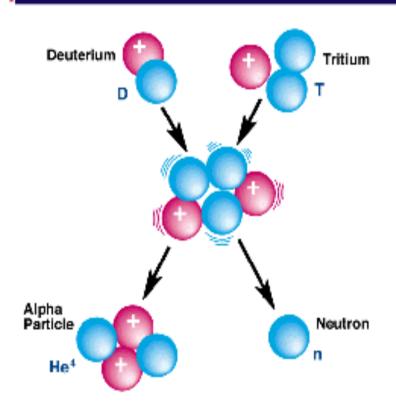
For D-T at T = 25 KeV, Lawson's criterium is:

$$n_e \tau > 1.5 \cdot 10^{14} \text{ s/cm}^3$$



Thermonuclear Fusion:

Deuterium-Tritium Fusion Reaction



Need to have high temperatures to overcome Coloumb repulsion

Need to have many fusion reactions to allow for energy gain, i.e. large number of particles and/or long confinement time.

 $n_{\rm p} \tau \approx 1.5 \ 10^{14} \ s \ cm^{-3}$

Lawson's criterium

"Triple product" $n_e \tau T \ge 8 \cdot 1014 \text{ s cm}^{-3} \text{ keV}$



Creating conditions for fusion:







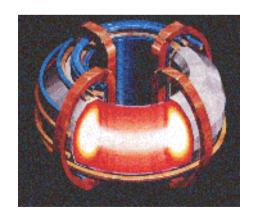
Heating Mechanisms:

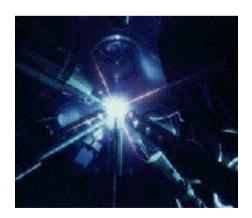
- * Compression (gravity)
- * Fusion Reactions (such as the p-p chain)

- * Electromagnetic Waves
- * Ohmic Heating (by electric currents)
- * Neutral Particle Beams (atomic hydrogen)
- * Fusion Reactions (D+T)

- * Compression (implosion driven by laser, or by X-rays from laser, or by ion beams)
- * Fusion Reactions (primarily D+T)









Creating conditions for fusion:



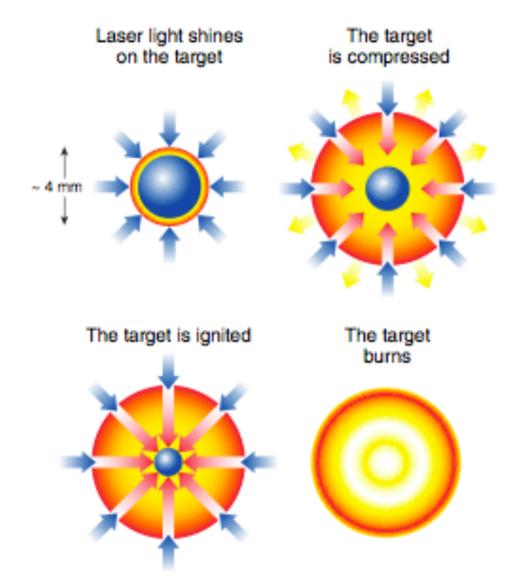
The secondary system (H-bomb) is ignited by the explosion of a "conventional" nuclear bomb

For controlled nuclear fusion:

- •Need to ignite small mass of fuel
- •Need to ignite with different tools!

Principio del Confinamento Inerziale

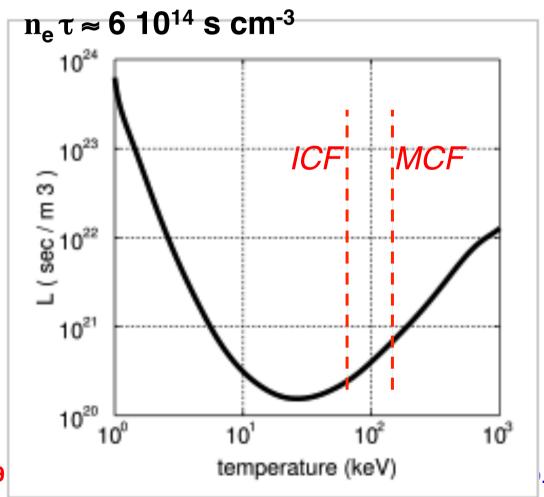




Lawson's criterium for inertial fusion



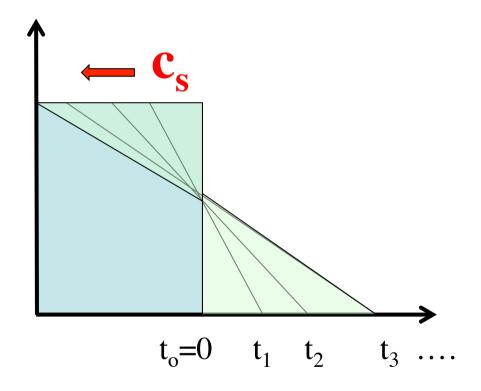
In ICFwe cannot reach 25 keV. Then





Isothermal expansion of a gas

Rarefaction (expansion) wave Self-similar model



$$n(x) = n_o \exp(-x/L)$$

$$L = c_s t$$

$$c_s = \left(\gamma Z k T / m\right)^{1/2}$$



Lawson's criterium for inertial fusion

$$n_e \tau \approx 6 \cdot 10^{14} \text{ s cm}^{-3}$$

Disassembly time determined by the fuels inertia

$$\tau < R/c_s \implies \text{we take } \tau = R/4c_s$$

ion sound velocity in a plasma

$$c_s = (\gamma Z k T_e / m_i)^{1/2} = 9.8 \quad 10^5 (\gamma Z T_e (eV) / \mu)^{1/2} cm / s$$

 $\approx 7 \quad 10^7 \quad cm / s \quad \text{(for T = 10 keV)}$

Notice:

in magnetic fusion the time τ expresses the *confinement of energy* In inertial fusion it refers to the *confinement of mass*





$$n_e \tau \approx 6 \ 10^{14} \ \text{s cm}^{-3}$$

Disassembly time determined by the fuels inertia

$$\tau = R/4c_s$$
 $c_s \approx 7 \cdot 10^7 \ cm/s \text{ (for T = 10 keV)}$

$$n_e = n_i = 2 \times \frac{\rho(g/cc)}{2.5} 6.022 \cdot 10^{23} cm^{-3} = 4.8\rho \cdot 10^{23} cm^{-3}$$

$$n_e \tau = 1.5 \cdot 10^{14} = (4.8\rho \cdot 10^{23})(R/4 \times 7.6 \cdot 10^7)$$

For typical ICF conditions

$$\Rightarrow \rho R = 1.5 \ 10^{14} \times 4 \times 7 \ 10^7 / 4.8 \ 10^{23} \approx 0.3 \ \text{g/cm}^2$$





$$\mathcal{E}_{\text{fus}} = \varepsilon_{DT} N_f = \frac{M_f}{2m_i} \, \Phi_B \varepsilon_{DT} = 3.4 \times 10^5 \Phi_B M_f$$

 $\Phi_{\rm R}$ "burned fraction" of the fuel

M_f mass of fuel (in g)

 $\mathcal{E}_{\mathrm{fus}}$ fusion energy

 ε_{DT} energy released in one reaction (17.6 MeV)

A few mg of DT are sufficient to produce an energy of several 100 MJ

To satisfy Lawson's criterion ($\rho R > 0.3 \text{ g/cm}^2$) at the density of solid D-T ($\rho_{\rm sol} = 0.25 \text{ g/cm}^3$) would require R = 1 cm. This implies a large explosion.

Burning fraction in ICF



We need to calculate the "burned fraction" of the fuel

$$R(t) = R_f - c_s t$$

$$t_{\text{max}} = R_f / c_s$$

$$N_{tot} = \frac{1}{2} n_i V_f$$

$$n_D = n_T = n_i / 2$$

$$V_f = \frac{4\pi}{3} R_f^3$$

$$N_f = n_D n_T \langle \sigma \mathbf{v} \rangle \int_0^{t_{\text{max}}} \frac{4\pi}{3} (R_f - c_s t)^3 dt =$$

$$= \frac{4\pi}{3} n_D n_T \langle \sigma \mathbf{v} \rangle \left[-\frac{1}{4c_s} (R_f - c_s t)^4 \right]_0^{t_{\text{max}}} = \frac{\pi}{3} n_D n_T \langle \sigma \mathbf{v} \rangle \frac{R_f^4}{c_s}$$

$$R(t) = R_f - c_s t \qquad t=0 R = R_f$$

$$t=t_{max} R=0$$

Burning fraction in ICF



$$N_{ion} = \rho_f V_f / m_i = n_i V_f$$
 total number of ions

$$N_{tot} = N_{ion} / 2$$

maximum number of reactions

$$N_f = \frac{\pi}{3} n_D n_T \langle \sigma \mathbf{v} \rangle \frac{R_f^4}{c_s} = \frac{1}{4m_i} \left[\left(\frac{4\pi}{3} R_f^3 \right) m_i n_i \right] \langle \sigma \mathbf{v} \rangle n_i \frac{R_f}{4c_s}$$

$$= \frac{1}{4m_i} \left[\rho_f V_f \right] \left\langle \sigma v \right\rangle n_i \frac{R_f}{4c_s} = \left\langle \sigma v \right\rangle N_{ion} \frac{R_f}{16c_s} = \left\langle \sigma v \right\rangle N_{tot} \frac{R_f}{8c_s}$$

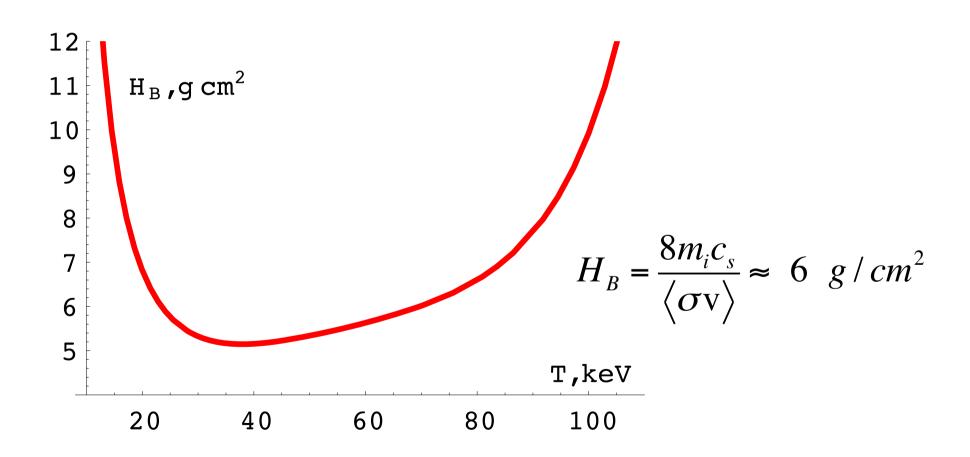
$$\Phi_B = \frac{N_f}{N_{tot}} = n_i \langle \sigma \mathbf{v} \rangle \frac{R_f}{8c_s} = n_i m_i \langle \sigma \mathbf{v} \rangle \frac{R_f}{8c_s m_i} = \frac{\rho_f R_f}{H_B}$$

$$H_B = \frac{8m_i c_s}{\langle \sigma v \rangle}$$

Burning fraction in ICF



H_B "combustion parameter" for a DT plasma fuel







The calculation didn't take into acount fuel consumption

$$n_f = n_{DO} - n_D$$

$$\frac{dn_f}{df} = -\frac{dn_D}{df} = n_D^2 \langle \sigma \mathbf{v} \rangle$$

$$n_D(t) = \frac{n_{DO}}{1 + n_{DO} \langle \sigma v \rangle t}$$

$$t_{\text{max}} = R_f / 4c_s$$

$$H_B = \frac{8m_i c_s}{\langle \sigma v \rangle} \approx 6 g/cm^2$$

$$n_{DO} = n_i / 2 = \rho_f / 2m_i$$





$$n_{D}(t_{\text{max}}) = \frac{n_{DO}}{1 + n_{DO} \langle \sigma v \rangle (R_{f} / 4c_{s})}$$

$$\Phi = \frac{n_{DO} - n_{D}(t_{\text{max}})}{n_{DO}} = 1 - \frac{n_{D}(t_{\text{max}})}{n_{DO}} = 1 - \frac{1}{1 + n_{DO} \langle \sigma v \rangle (R_{f} / 4c_{s})}$$

$$= 1 - \frac{1}{1_{i} + (\rho / 2m_{i}) \langle \sigma v \rangle (R_{f} / 4c_{s})} = 1 - \frac{1}{1 + \rho R_{f} / H_{B}}$$

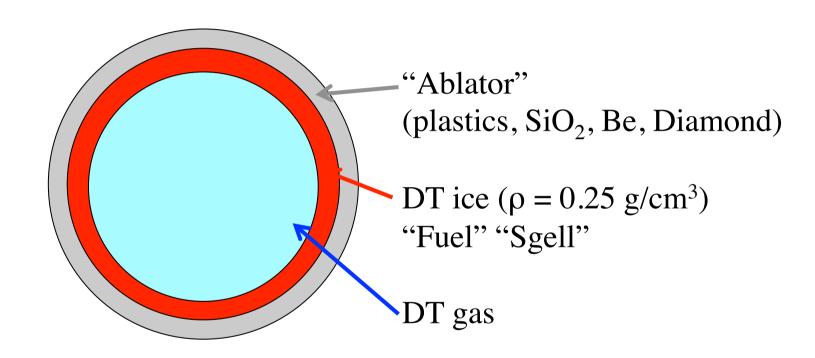
$$= \frac{(1 + \rho R_{f} / H_{B}) - 1}{1 + \rho R_{f} / H_{B}} = \frac{\rho R_{f} / H_{B}}{1 + \rho R_{f} / H_{B}} = \frac{\rho R_{f}}{H_{B} + \rho R_{f}} \approx \frac{\rho R_{f}}{\rho R_{f} + 6}$$

Conventionally we take $\rho R \approx 3 \text{ g/cm}^2$ that is 33% burned fuel

$$H_B = \frac{8m_i c_s}{\langle \sigma v \rangle}$$



ICF typical targets

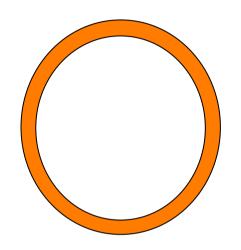


External radius ≈ mm

DT mass ≈ mg



ICF typical targets



INITIAL CONDITIONS

CRYOGENIC SHELL

- $R_{in} \approx 2 \text{ mm}$, $R_{in} / \Delta r \approx 60 (\Delta r \approx 33 \mu \text{m})$
- V ≈ $4 \pi R_{in}^{2} \Delta r \approx 1.6 \cdot 10^{-3} cm^{3}$
- $\rho_{in} \approx 2.5 \times~0.2~g/~cm^3,~M \approx 0.85~mg$

FINAL CONDITIONS

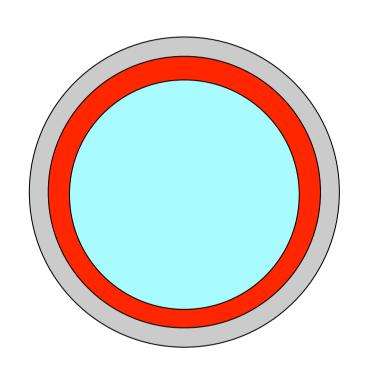
- ρ_{fin} / ρ_{in} ≈ 1000
- ρ_{fin} ≈ 500 g/ cm³
- $R_{fin} \approx 74 \ \mu m$
- ρ_{fin} R_{fin} \approx 3.7 g/ cm²

Hot spot

Compressed fuel



Some orders of magnitude

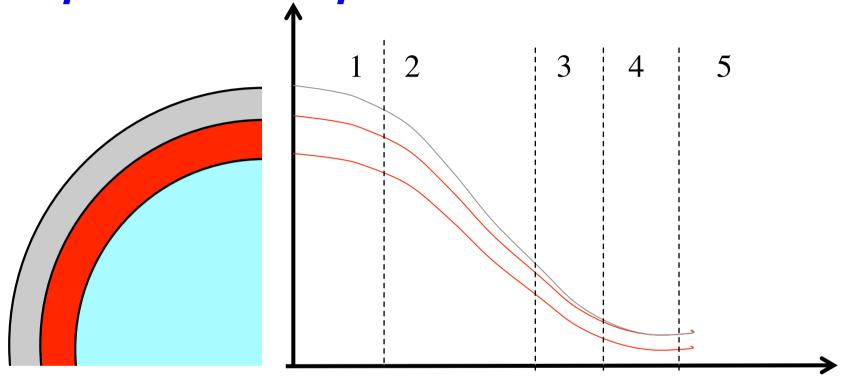


```
If
t_{laser} \approx t_{implosion} \approx 10 \text{ ns}
I_{L} \approx 3 \ 10^{14} \text{ W/cm}^{2}
R \approx 2 \text{mm}
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Then S=0.5 \text{ cm}2 E_{laser} = I_L t_{laser} S=1.5 \text{ MJ} V_{implosion} = 2 \text{ mm/}10 \text{ ns} =200 \ \mu\text{m/ns} = 200 \text{ km/s}
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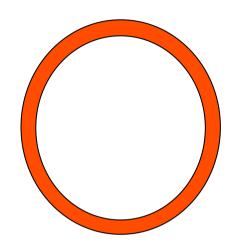
Space-time plot

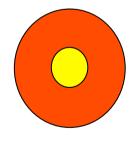


- 1 ablation and acceleration
- 2 implosion (almost constant velocity)
- 3 deceleration
- 4 stagnation (creation of hot spot)
- 5 explosion



Why do we need a hot spot?





INITIAL CONDITIONS

CRYOGENIC SHELL

- $R_{in} \approx 2 \text{ mm}$, $R_{in} / \Delta r \approx 60 (\Delta r \approx 33 \mu \text{m})$
- V \approx 4 π R_{in}² Δ r \approx 1.6 10⁻³ cm³
- $\rho_{\text{in}} \approx 2.5 \times 0.2 \text{ g/cm}^3$, M $\approx 0.85 \text{ mg}$

FINAL CONDITIONS

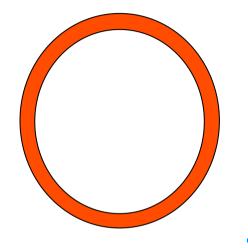
- ρ_{fin} / ρ_{in} \approx 1000
- $\rho_{\text{fin}} \approx 500 \text{ g/cm}^3$ $R_{\text{fin}} \approx 74 \text{ }\mu\text{m}$
- ρ_{fin} R_{fin} \approx 3.7 g/ cm²

Total number of ions $N_{DT} \approx 2 \cdot 10^{20}$

If $T_{fin} \approx 10 \text{ keV}$, total thermal energy in fuel $E \approx 2 (3/2 \text{ N}_{DT} \text{ T}) \approx 1 \text{ MJ } !!$

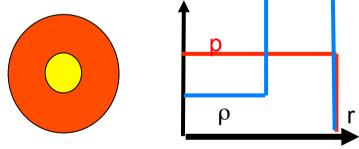


"Isobaric" approach to ICF



Stagnation is reached as the pressure of the internal fuel increases and gradually slows the shell down.

At stagnation P_{shell} ≈ P_{central spot}



Kinetic Energy of (remaining) imploding shell is converted in:

- Compression of the fuel in the shell
- Heating of the central has

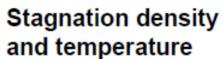
Produces an Isobaric fuel assembly

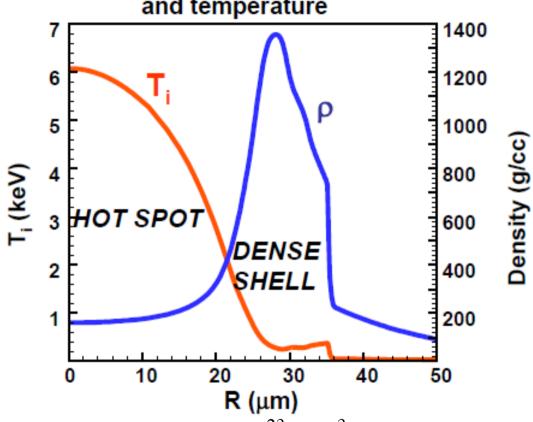
Conversion efficiency from laser ligth to kinetic energy of the shell is extremely low $\approx 5 \%$

Need for High Aspect ratio target and high implosion velocities V ~400 km/s



NIF-like target (1 MJ)





$$n_e = n_i = 4.8 \rho \ 10^{23} cm^{-3}$$

$$P(Bar) = 1.8 \ 10^{-18} \ n_{TOT}(cm^{-3}) \ T_e(eV)$$

SHELL

$$-\rho_{fin} R_{fin} \approx 10^3 \text{ g/cm}^3 \times 40 \mu \text{m}$$

= 4 g/ cm²
-n_i ≈ 5 10²⁶ cm⁻³

HOT SPOT

$$-\rho_{fin} R_{fin} \approx 130 \text{ g/cm}^3 \times 22 \mu \text{m}$$

= 0.3 g/cm²

PRESSURE

-P≈ 300 GBar



Spherical geometry

Notice:

Ablation Pressure P ≈ 50 MBar

Pressure at stagnation P ≈ 50 Gbar

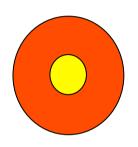
Amplification of a factor \times 1000 due to convergence.

Spherical geometry is essential for ignition



Energy balance in the fuel

The compressed shell is a dense degenerate plasma



$$T_F = (3\pi^2 n)^{2/3} \hbar^2 / 2m,$$

 $T_F \simeq 14 \,\rho^{2/3} \,\text{eV},$

$$p_F = \frac{2}{5}n_e T_F.$$
 $p_F = A_F \rho^{5/3} = 2.16 \, \rho^{5/3} \text{Mbar}.$

"entropy parameter"

$$lpha=p/p_F,$$

$$p=lpha\,{\rm A_F}\,
ho^{5/3}$$

The hot spot is a hot classical plasma

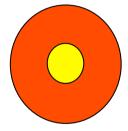


Energy balance in the fuel

The compressed shell is a dense degenerate plasma

$$\mathcal{E}_{\text{compr}} = \int_{V_f}^{V_0} p \, dV = \alpha M_f \int_{\rho_0}^{\rho_f} \frac{p_F}{\rho^2} \, d\rho \simeq \frac{3}{2} \alpha M_f A_F \rho_f^{2/3} = 0.35 \, \alpha \, M_f \rho_f^{2/3} \, \text{MJ}.$$

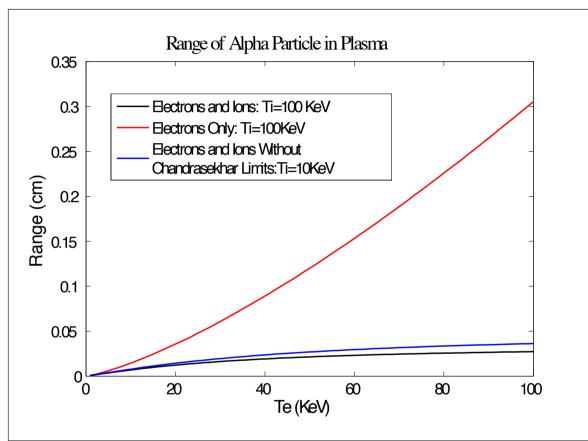
The hot spot is a hot classical plasma



$$\mathcal{E}_{\text{chauf}} = \frac{3}{2} (n_e + n_D + n_T) T_h V_h = 3 n_e T_h V_h = 110 M_h T_{\text{keV}} \text{ MJ}.$$



α particle range in ICF plasma



$$R_a = 0.107 \frac{T_e^{3/2}}{\rho \ln (\Lambda)} [cm]$$

Assuming 50/50 DT plasma (T_e in keV, ρ in g/cm³). The plot of the range is shown in figure for $\rho = 50$ g/cm³ and $T_i = 100$ keV

To get $R_a \approx 10 \, \mu \text{m}$, one needs $\rho \approx 500 \, \text{g/cm}^3$

In this case α particles are confined within the plasma and contribute to its heating

Lessons from History ...



Laser Compression of Matter to Super-High Densities: Thermonuclear (CTR) Applications
JOHN NUCKOLLS, LOWELL WOOD, ALBERT THIESSEN & GEORGE ZIMMERMAN
University of California Lawrence Livermore Laboratory

Nature **239**, 139 - 142 (15 September **1972**)

Hydrogen may be compressed to more than 10,000 times liquid density by an implosion system energized by a high energy laser. This scheme makes possible efficient thermonuclear burn of small pellets of heavy hydrogen isotopes, and makes feasible fusion power reactors using practical lasers.

Early predictions for laser ignition were ~1kJ [Nuckolls]

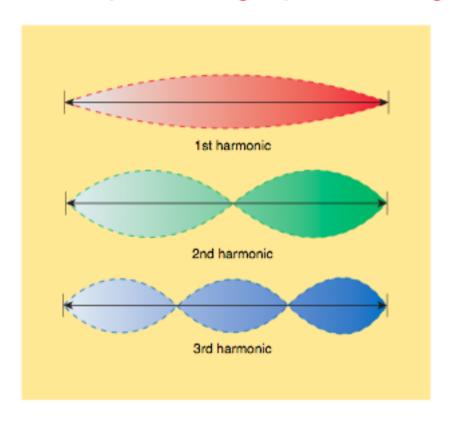
- What was wrong?
 - Very strong sensitivity to implosion velocity
 - Optimistic assessment of hydro instability growth
 - Assumption of high coupling efficiency at high laser irradiance

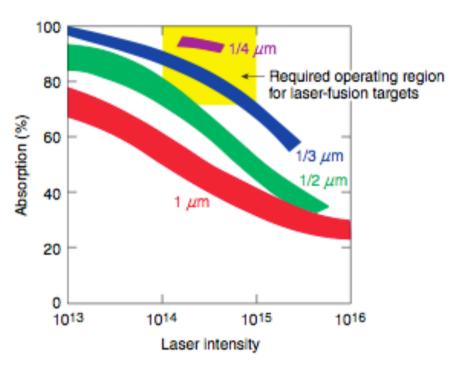
Note: estimates for high gain have remained ~ constant (at ~MJ), as much weaker dependence on implosion velocity





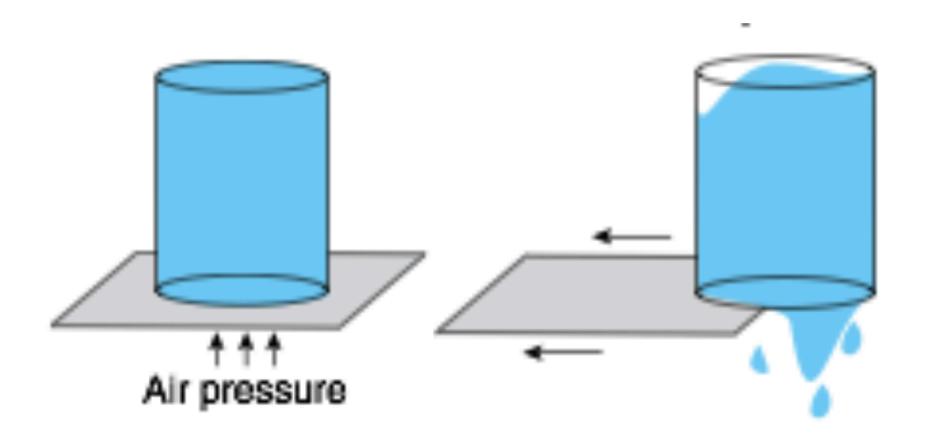
Laser Matter Interaction prohibits large intensities (low absorption, high preheating)





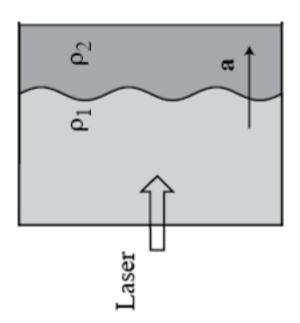


Rayleigh-Taylor Instability:





Rayleigh-Taylor Instability:



Use Fourier expansion (we write a sinusoidal perturbation).

Here
$$A(x,t) = A_0 \cos(kx - \omega t)$$

The equation of dynamics bring to write a dispersion relation $\varepsilon(\omega,t) = 0$

If ω is real stable behaviour (wave which is eventually damped and dies away)

If $\omega = i\gamma$ is imaginary then the perturbation grows like $A(x,t) = A_0 \cos(kx) e^{\gamma t}$



Rayleigh-Taylor Instability:

"Classical" growth rate (linear phase)

$$\gamma = \sqrt{Akg}$$

$$A = \sqrt{\frac{\rho_2 - \rho_1}{\rho_2 - \rho_1}}$$

γ (ns-1)

Atwood number (instable if $\rho_1 < \rho_2$)

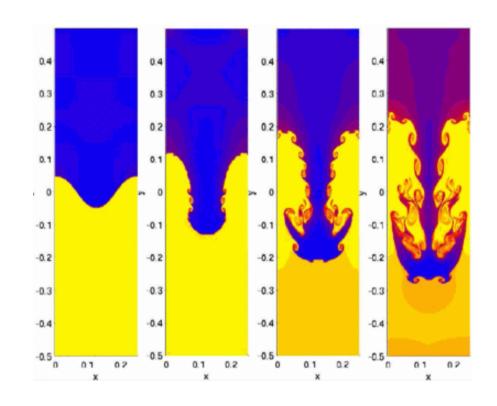
$$\gamma = \sqrt{\frac{Akg}{1 + kL}} - \beta k v_{abl}$$

Modified "Takabe" expression. L plasma density gradient

Rayleigh-Taylor Instability



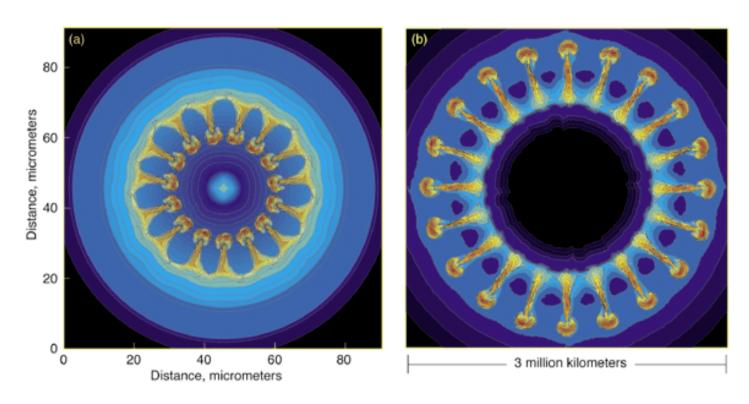
- Major instability: heavy material pushes on low density one
- Will always occur since driver is never 100% symmetric
- The Rayleigh-Taylor instability always grows



Energy must be delivered as sysmmetric as possible!



Rayleigh-Taylor Instability – spherical implosions /



Striking similarities exist between hydrodynamic instabilities in (a) inertial confinement fusion capsule implosions and (b) core-collapse supernova explosions. [Image (a) is from Sakagami and Nishihara, *Physics of Fluids B* 2, 2715 (1990); image (b) is from Hachisu et al., *Astrophysical Journal* 368, L27 (1991).]

Energy must be delivered as sysmmetric as possible!



Good news:

Experiments with GEKKO XII laser, Institute of Laser Engineering, University of Osaka, Japan

- Experimental demonstration of compression of DT up to 600 × solid density (Azechi et al., Las. Part. Beams, 1991)
- WE ARE ABLE TO OBTAIN DENSITIES DIRECTLY RELEVANT FOR ICF!

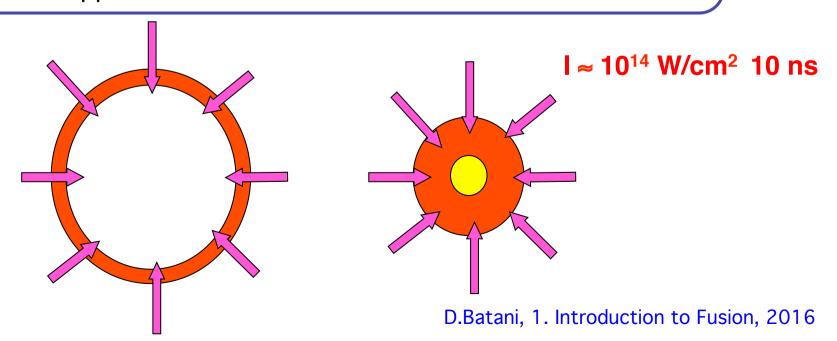
However: number of neutrons much smaller than expected: The central hot spot was not generated



ICF: direct drive approach

Lawson's Criterium (buring criterum) for ignition (D-T): ρ R > 3 gcm⁻² and T ≈ 5 - 10 keV

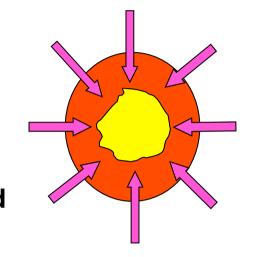
- synchronized laser pulses with spherical irradiation symmetry
- shock wave compression (up to 1000 × solid density)
- ignition from a central hot spark produced by shock coalescence
- isobaric approach





Problems of classical scheme:

Non uniformities in laser irradiation or in target bring to:



- Mixing of fuel and wall, higher Z*, increased emission and cooling
- The central hot spot is not generated

$$\frac{\Delta R_{fin}}{R_{fin}} \approx \frac{\Delta v}{R_{fin}} t_{imp} \approx \frac{R_{in}}{R_{fin}} \frac{\Delta v}{v_{imp}} \approx \frac{R_{in}}{R_{fin}} \frac{\Delta I}{I}$$

Then:

$$\frac{\Delta R_{fin}}{R_{fin}} \approx 50\% \Rightarrow \frac{\Delta I}{I} \approx 1\%$$

How to relax uniformity constraints for ICF?



- **USE INDIRECT DRIVE**
- >USE OPTICAL SMOOTHING
- >USE OF "FOAM BUFFERED TARGETS"

Willi et al., Phys. Rev.Lett., 1995

> SEPARATE COMPRESSION AND IGNITION **Fast Ignition**

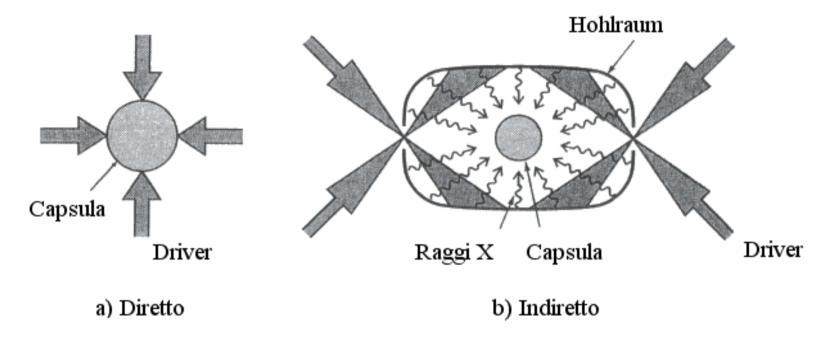
Tabak et al., Phys. Plasmas, 1994

Shock Ignition

R. Betti, et al. Phys. Rev. Lett., 2007



Inertial confinement: direct vs. indirect drive



Direct: higher efficiency, more problems with uniformity

Indirect: better uniformity but reduction of efficiency

In both case you need MJ-class laser systems



Targets

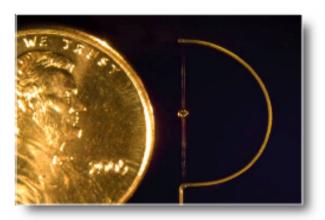


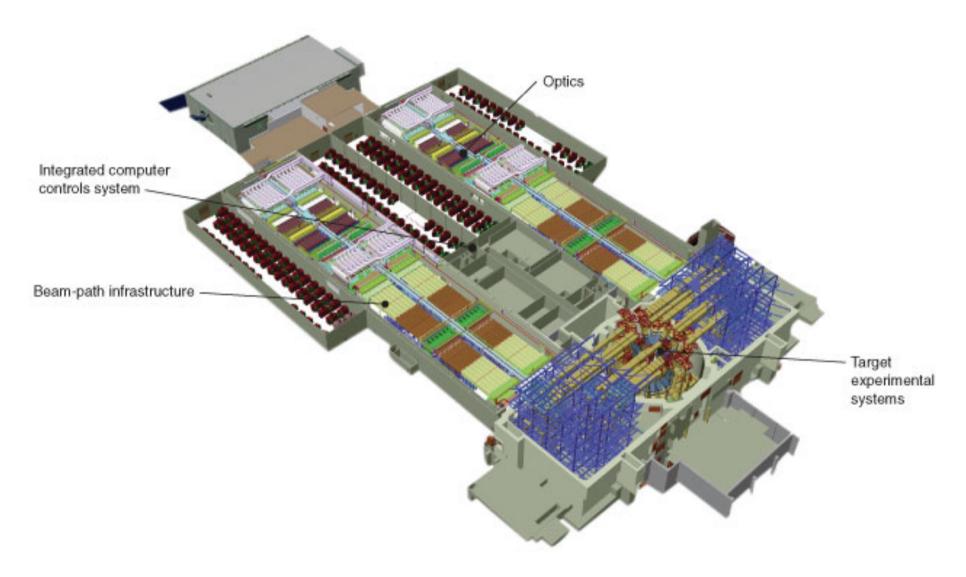
Figure 3.3: The laser-fusion targets used on OMEGA experiments are typically ~1 mm in diameter and are suspended by spider silks.



Figure 3.7: A typical indirect-drive target.

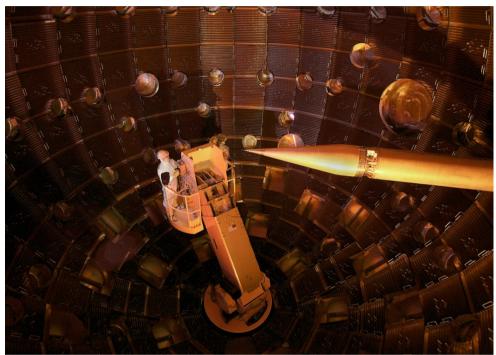
National Ignition Facility (NIF) - layout

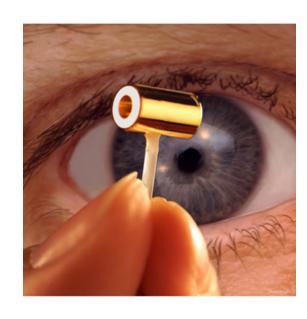




NIF



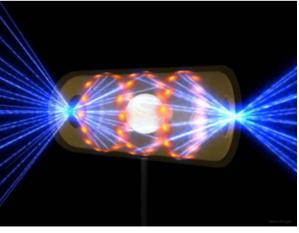




Indirect drive

- Chamber
- Target holder
- Hohlraum
- Pellet



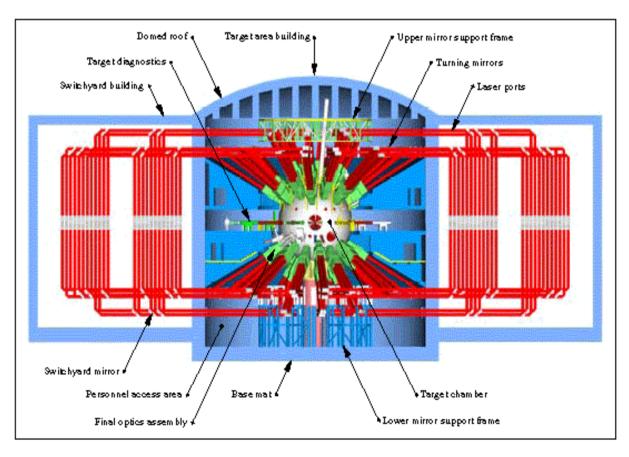


D.Batani, 1. Introduction to Fusion, 2016



Lasers - NIF and Megajoule:

Controlled Thermonuclear Fusion

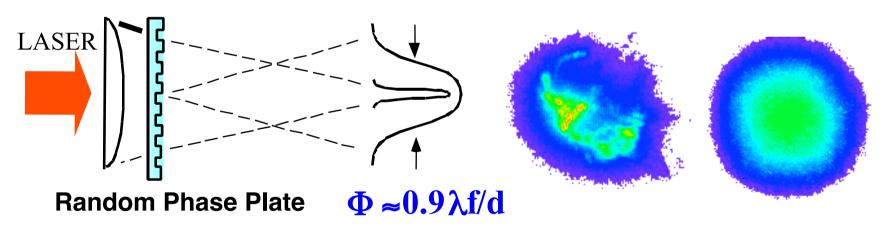


Nd:glass **2 MJ** 10 ns 200 beams



Optical smoothing techniques:

Optical smoothing techniques (RPP, ISI, SSD..) introduced (80's, 90's) to produce "gaussian" beam profiles with small scale modulations



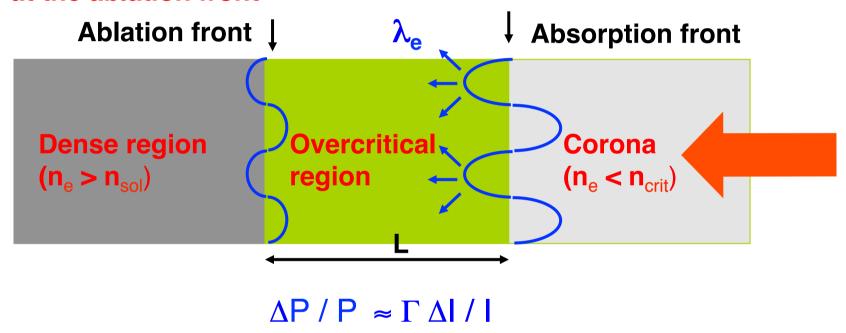
2D square elements with **0** or π dephasing (*Kato*, *PRL*, 1984)

Small scale modulations are rapidly washed out by thermal smoothing



Thermal smoothing:

Laser is absorbed at the absorption front but pressure is applied at the ablation front



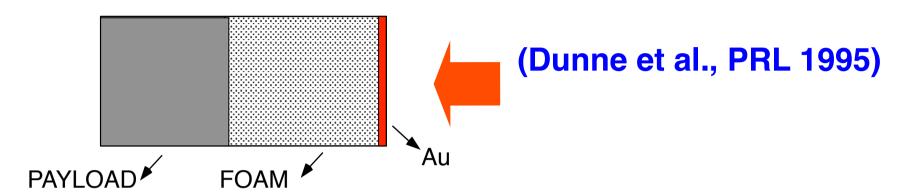
Non-uniformities present at the laser absorption surface are reduced at the ablation surface by

$$\Gamma \approx \exp \left[-\alpha L \left(2\pi / \lambda_{perp} \right) \right]$$

L "stand-off distance"



Foam buffered targets:



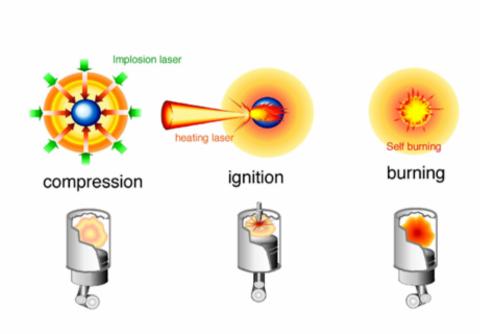
CREATE A SUPERCRITICAL AND SUFFICIENTLY WIDE PLASMA
IN FRONT OF THE TARGET

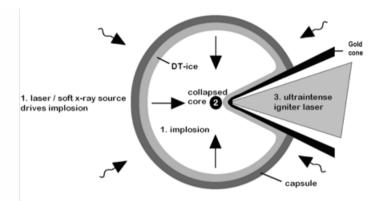
OPEN QUESTIONS:

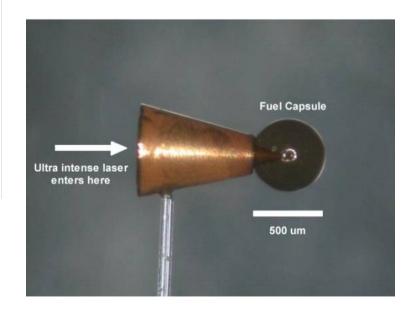
- Optimal parameters of foam and converter
- What λ_{perp} ?
- Effects on hydrodynamics? Shock heating of DT?



Inertial Fusion- Fast Ignition







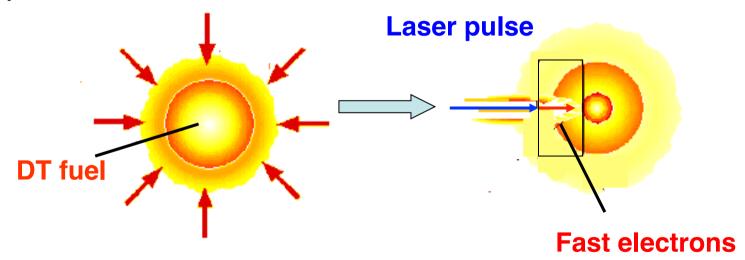
D.Batani, 1. Introduction to Fusion, 2016

The concept of 'fast ignition'



>1: "normal" compression with ns laser beams (we are able to compress!)

▶2: a CPA laser creates a beam of relativistic electrons (lateral hot spot)

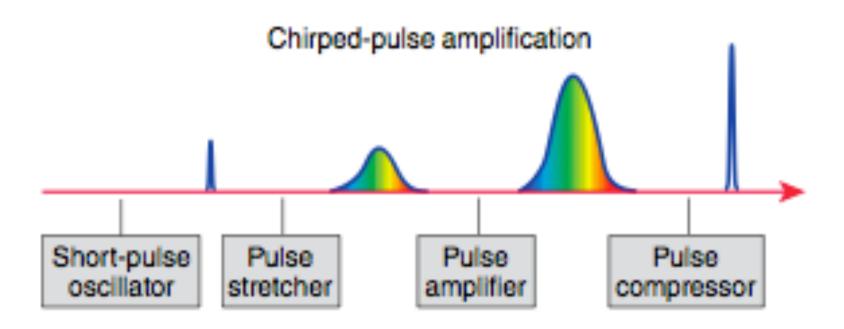


Typical parameters: E ≈ 10 kJ, Δt ≈ 10 ps, R ≈ 10μm, E_{fast} ≈ 1 MeV



High-intensity lasers (high energy – short pulse)

Laser-plasma interaction in the "relativistic regime"





Evolution of laser performance:

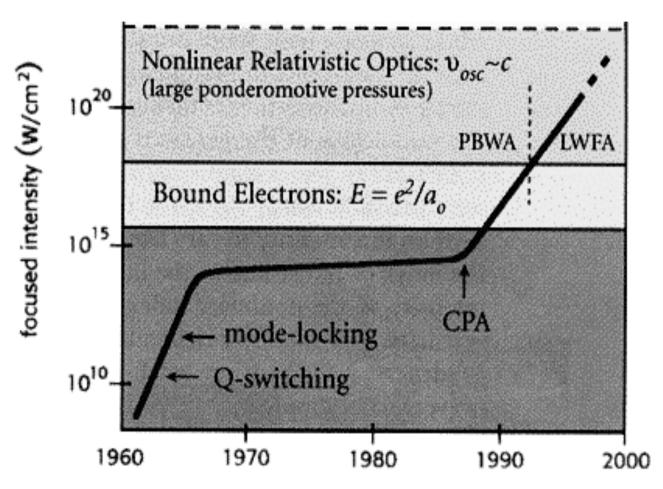
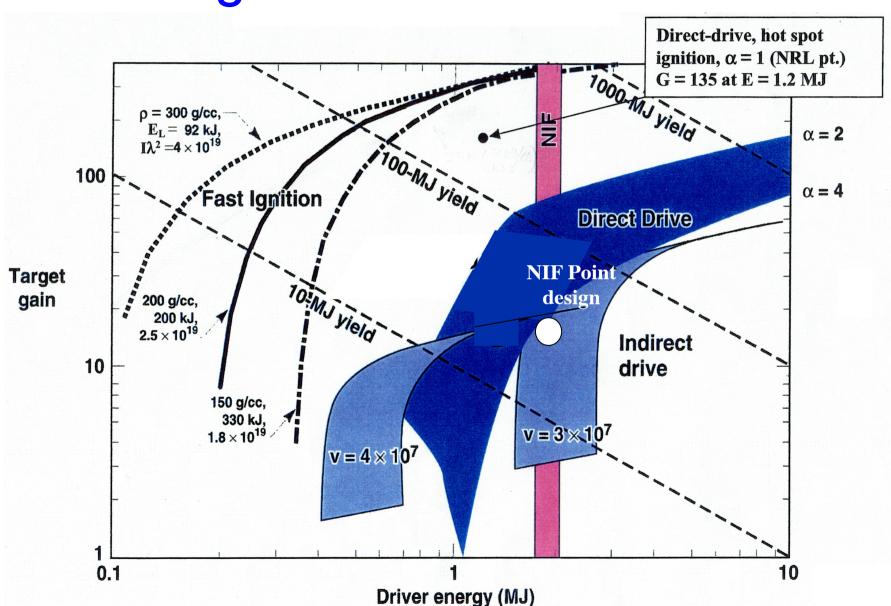


Figure 1. Laser intensity as a function of year, showing the impact of the CPA concept and the different thresholds of physical regimes. The sharp increase in intensity since the advent of CPA is comparable to the sharp increase after the invention of the laser in the 1960's.



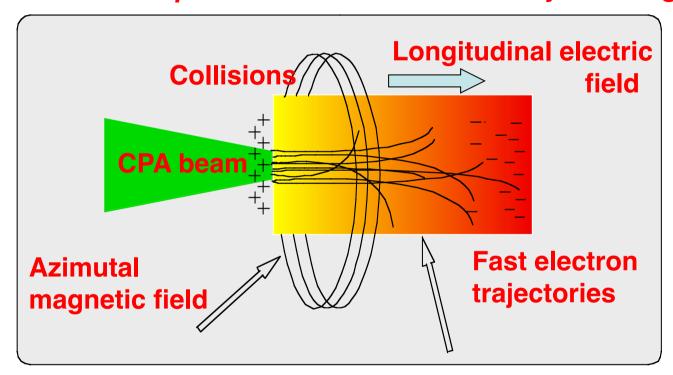
Advantage of FI





Study of propagation:

Propagation of fast electrons in matter between n_c and 100 n_c over 200 - 300 μ m is critical for the feasibility of fast ignition



- Collisional Effects (Stopping Power)
 NEW PHYSICAL PROBLEM:
- Electric field effects
- Magnetic fields effects

NEW PHYSICAL PROBLEM: VERY DIFFERENT FROM JUST BETHE-BLOCH

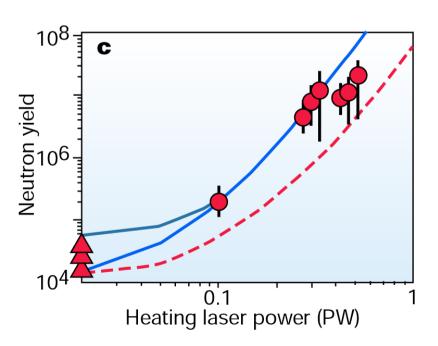


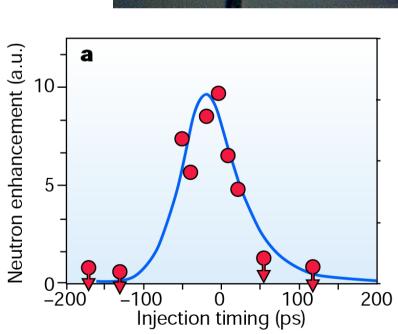
Use of cone targets

"Cone guided" targets were tested at the ILE in Osaka

- The laser can interact in regions at higher densities
- Increment in neutron yield from compressed fusion targets

R. Kodama et al., Nature 418, 933 (2002)





Physical issues



- Cones do not seem viable for future reactors (same problems of ID targets)
- The presence of the cone may prevent a sufficiently uniform compression and produce high-Z pollution
- How to control electron beam energy?
- Is there any tools to control electron beam divergence?

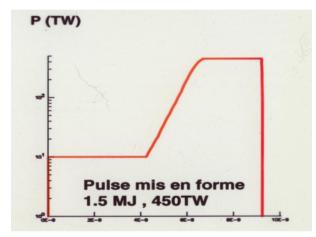
Although a very promising idea, FI is in a "premature" stage for what concerns technological developments (we need a 100 kJ - 10 ps laser beam)

[the main issue here is scalability of collective effects...]

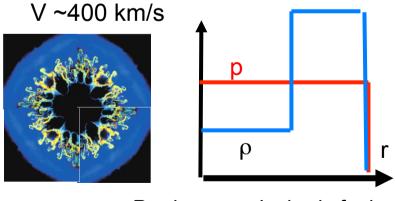
A final laser spike launches a converging CELIA shock



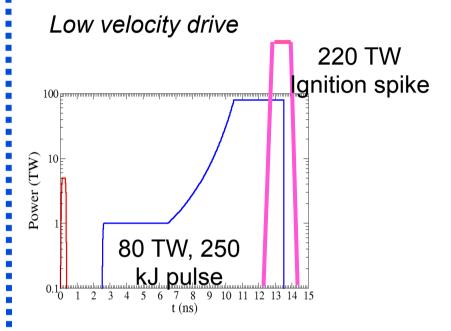
Conventionnal direct drive 450 TW, 1.5MJ pulse



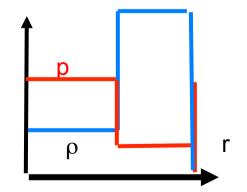
High Aspect ratio target



Produces an Isobaric fuel assembly



Low AR V~240 km/s



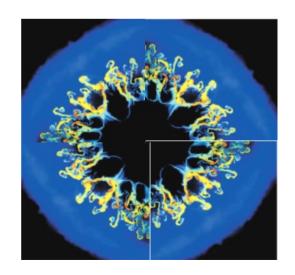
Fuel assembly is non isobaric

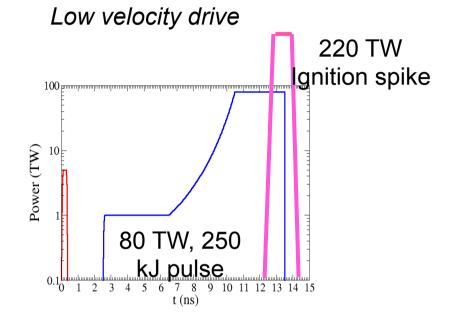
D.Batani, 1. Introduction to Fusion, 2016

Shock ignition is less sensitive to hydro instabilities

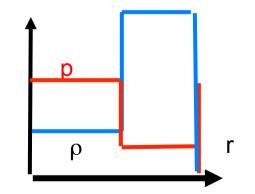


In SI, you do not create the hot spot with the "main" compression beam. Hence you do not need such highh implosion velocity. Hence you can implode a more massive thicker shell which does not break due to RT









Fuel assembly is non isobaric

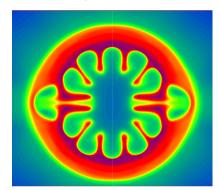
D.Batani, 1. Introduction to Fusion, 2016

The Ignition shock mitigates RT growth at stagnation



HiPER target at time of maximum ρ R (1D)

180 kJ 48 beams



HiPER target at ignition time

